

## Introduction to String Theory

Humboldt-Universität zu Berlin  
 Dr. Emanuel Malek

### Exercise Sheet 1

**1** Consider a coordinate change

$$x^\mu \longrightarrow x'^\mu. \quad (1.1)$$

Recall that the components of tensors transform in a well-defined way under the coordinate change (1.1). For example, the components of a (1, 1) tensor  $T^\mu{}_\nu$  transform under (1.1) as

$$T^\mu{}_\nu \longrightarrow T'^\mu{}_\nu = \frac{\partial x'^\mu}{\partial x^\rho} \frac{\partial x^\sigma}{\partial x'^\nu} T^\rho{}_\sigma. \quad (1.2)$$

Similarly, the components of a (1, 1) tensor density  $\tilde{T}^\mu{}_\nu$  of weight  $w$  transforms under (1.1) as

$$\tilde{T}^\mu{}_\nu \longrightarrow \tilde{T}'^\mu{}_\nu = |J|^w \frac{\partial x'^\mu}{\partial x^\rho} \frac{\partial x^\sigma}{\partial x'^\nu} T^\rho{}_\sigma, \quad (1.3)$$

where  $J = \det \left( \frac{\partial x'^\mu}{\partial x'^\nu} \right)$  is the Jacobian of the coordinate change (1.1).

(a) Given a (2, 0) tensor  $S_{\mu\nu}$ , show that  $(\det S_{\mu\nu})^{1/2}$  is a scalar density of weight 1.

(b) Consider now tensor and tensor density fields, i.e.  $T^\mu{}_\nu(x)$ , etc. The transformation (1.1) can be expanded infinitesimally as

$$x'^\mu = x^\mu - \epsilon^\mu(x) + \dots \quad (1.4)$$

Show the following infinitesimal variations for a scalar fields  $\Phi(x)$ , Lorentzian metric  $g_{\mu\nu}(x)$  and its associated density  $(-\det g(x))^{1/2}$ :

$$\begin{aligned} \delta\Phi &= \epsilon^\mu \partial_\mu \Phi, \\ \delta g_{\mu\nu} &= \epsilon^\gamma \partial_\gamma g_{\mu\nu} + (\partial_\mu \epsilon^\gamma) g_{\gamma\nu} + (\partial_\nu \epsilon^\gamma) g_{\mu\gamma} = \nabla_\mu \epsilon_\nu + \nabla_\nu \epsilon_\mu, \\ \delta(-\det g)^{1/2} &= \partial_\gamma (\epsilon^\gamma (-\det g)^{1/2}). \end{aligned} \quad (1.5)$$

The second equality of the second line of (1.5) uses the torsion-free metric-compatible connection  $\nabla_\mu$ .

*Hint:* A scalar field is defined via  $\Phi'(x') = \Phi(x)$  and  $\delta\Phi$  in (1.5) is defined as  $\delta\Phi(x) \equiv \Phi'(x) - \Phi(x)$  and similarly for the other fields in (1.5).

**2** Consider a submanifold  $H$  of  $M$  defined by the embedding  $\mathcal{Y} : H \hookrightarrow M$ . A Riemannian metric  $g$  on  $M$  then induces a pull-back metric  $h = \mathcal{Y}^*g$  on  $H$ , explicitly given by

$$ds^2 = d\mathcal{Y}^a d\mathcal{Y}^b g_{ab}, \quad (2.1)$$

where  $a, b = 1, \dots, \dim M$ , and  $d\mathcal{Y}^a$  denotes the exterior derivative of  $\mathcal{Y}^a$  on  $H$ . In local coordinates  $\sigma^\alpha$ ,  $\alpha = 1, \dots, \dim H$ , on  $H$ , these are explicitly given by

$$d\mathcal{Y}^a = \frac{\partial \mathcal{Y}^a}{\partial \sigma^\alpha} d\sigma^\alpha. \quad (2.2)$$

Consider the unit-radius  $n$ -sphere  $S^n \subset \mathbb{R}^{n+1}$  with the embedding

$$\mathcal{Y} : S^n \hookrightarrow \mathbb{R}^{n+1}, \quad (2.3)$$

and where  $\mathbb{R}^{n+1}$  is endowed with the flat Riemannian metric  $g_{ab} = \delta_{ab}$ , such that the embedding satisfies  $\mathcal{Y}^a \mathcal{Y}^b \delta_{ab} = 1$ .

(a) Compute the pull-back metric (2.1) for  $S^2$  using spherical coordinates

$$\mathcal{Y}^a = (\cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta), \quad (2.4)$$

with  $\phi \in [0, 2\pi]$  and  $\theta \in [0, \pi]$ .

(b) Compute the pull-back metric (2.1) for  $S^n$  using local coordinates

$$\mathcal{Y}^a = \left( y^i, \sqrt{1 - y^i \delta_{ij} y^j} \right), \quad (2.5)$$

where  $i = 1, \dots, n$ .

(c) Compute the pull-back metric (2.1) for  $\mathcal{Y} : H^{p,n-p} \hookrightarrow \mathbb{R}^{n+1}$ , where  $\mathbb{R}^{n+1}$  is endowed with the flat pseudo-Riemannian metric

$$\eta_{ab} = \text{diag}(\underbrace{-1, \dots, -1}_{n-p}, \underbrace{+1, \dots, +1}_{p+1}), \quad (2.6)$$

and the embedding  $\mathcal{Y}^a$  satisfies  $\mathcal{Y}^a \mathcal{Y}^b \eta_{ab} = 1$ .

*Hint:* Use local coordinates on  $H^{p,n-p}$  such that

$$\mathcal{Y}^a = \left( y^i, \sqrt{1 - y^i \eta_{ij} y^j} \right), \quad (2.7)$$

where  $i = 1, \dots, n$  and  $\eta_{ij}$  denotes the first  $n \times n$  block of  $\eta_{ab}$  in (2.6).

**3** Consider the action for a field  $\phi(x)$  in a  $d$ -dimensional spacetime with Minkowski metric  $\eta_{\mu\nu} = (-1, +1, \dots, +1)$ . The action is given in terms of the Lagrangian density by

$$S[\phi] = \int d^d x \mathcal{L}(\phi(x), \partial_\mu \phi(x)). \quad (3.1)$$

(a) Use integration by parts and assume that you can neglect boundary terms, to show that the action is minimised  $\delta S = 0$  for field configurations satisfying the Euler-Lagrange equations

$$\partial_\mu \left( \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \right) = \frac{\partial \mathcal{L}}{\partial \phi}. \quad (3.2)$$

(b) Consider the Klein-Gordon action for a massive scalar field  $\phi$  in  $d$  dimensions with Minkowski metric

$$S = -\frac{1}{2} \int d^d x (\partial_\mu \phi \partial^\mu \phi + m^2 \phi^2) \quad (3.3)$$

Use the action (3.3) to compute the Euler-Lagrange equations for the scalar field  $\phi$ .

(c) Show that the action (3.3) is invariant under translations and use Noether's Theorem to compute the stress-energy tensor associated with this translation invariance.

(d) Compute the stress-energy tensor in the following alternative way. Consider a position-dependent translation

$$x^\mu \longrightarrow x^\mu + \epsilon^\mu(x). \quad (3.4)$$

The variation of the action (3.3) takes the form

$$\delta S = \int d^d x T^{\mu\nu} \partial_\mu \epsilon_\nu. \quad (3.5)$$

Compute  $T^{\mu\nu}$  from (3.5).

(e) Convince yourself that the variation of any translation-invariant action under a position-dependent translation (3.4) must take the form (3.5). Moreover, use (3.5) to argue that  $T_{\mu\nu}$  must be conserved when the equations of motion hold.

*Hint:* Consider what happens when  $\epsilon$  is constant and what it means for the action to be extremised.

(f) Compute the stress-energy tensor in the following alternative way. Minimally couple the Klein-Gordon action to a dynamical metric  $g_{\mu\nu}$  and compute the stress-energy tensor via

$$T_{\mu\nu} = -\frac{2}{\sqrt{-g}} \left. \frac{\delta S}{\delta g^{\mu\nu}} \right|_{g=\eta}, \quad (3.6)$$

where  $|_{g=\eta}$  denotes that we are evaluating with the metric  $g$  set to the Minkowski metric  $\eta$ .

(g) Think about the translation-dependent shift (3.4) as a diffeomorphism to show that the stress-energy tensors computed in (3.5) and (3.6) must be the same for any action.

4 Consider the following 1-dimensional action

$$S = \frac{1}{2} \int_{\mathcal{P}} d\tau \left( e^{-1} \dot{X}^\mu \dot{X}^\nu \eta_{\mu\nu} - e m^2 \right), \quad (4.1)$$

where  $X^\mu(\tau)$ ,  $e(\tau)$  are dynamical fields,  $\dot{X}^\mu = \frac{dX^\mu}{d\tau}$ ,  $m$  is a constant and  $\eta_{\mu\nu}$  is the Minkowski metric with  $\mu = 0, \dots, n$ .

(a) Given that  $X^\mu(\tau)$  are scalar fields, how does  $e$  have to transform under reparameterisations

$$\tau \longrightarrow \tilde{\tau}(\tau), \quad (4.2)$$

such that the action (4.1) is invariant?

(b) Show that the equations of motion for  $e$  imply

$$e^2 m^2 + \dot{X}^\mu \dot{X}^\nu \eta_{\mu\nu} = 0. \quad (4.3)$$

Plug (4.3) back into the action (4.1) to show that the action is equivalent to

$$S = -m \int_{\mathcal{P}} d\tau \sqrt{-\dot{X}^\mu \dot{X}^\nu \eta_{\mu\nu}}. \quad (4.4)$$

This is the world-line action of a point-particle in Minkowski space.

(c) Show that the action (4.4) is the same as

$$S = -m \int_{\mathcal{P}} d\tau \sqrt{-\det(X^*\eta)}, \quad (4.5)$$

where  $X^*\eta$  is the pull-back of the Minkowski metric  $\eta$  to the worldline of the particle,  $\mathcal{P}$ , embedded in Minkowski space:  $X : \mathcal{P} \hookrightarrow \mathbb{R}^{1,n-1}$ .

(d) Compute the equation of motion for the action (4.4). Write the equation of motion in terms of the canonical momentum for  $X^\mu$  and interpret this equation.

(e) Consider a charged point-particle in Minkowski space coupled to an electromagnetic field

$$S = -m \int_{\mathcal{P}} d\tau \sqrt{-\det(X^*\eta)} + q \int_{\mathcal{P}} d\tau X^* A, \quad (4.6)$$

where  $X^*A$  denotes the pull-back of the gauge potential  $A_\mu$  to the worldline of the particle. Convince yourself that the action is explicitly given by

$$S = -m \int_{\mathcal{P}} d\tau \sqrt{-\frac{dX^\mu}{d\tau} \frac{dX^\nu}{d\tau} \eta_{\mu\nu}} + q \int_{\mathcal{P}} d\tau A_\mu(X(\tau)) \frac{dX^\mu}{d\tau}. \quad (4.7)$$

Compute and interpret the equation of motion for the action (4.6).

*Hint:* Recall the definition of the electromagnetic field strength tensor  $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ .

(f) Consider a neutral point-particle moving in a curved space-time  $(M, g)$ , i.e.

$$S = -m \int_{\mathcal{P}} d\tau \sqrt{-\det(X^*g)}, \quad (4.8)$$

where  $X^*g$  is now the pull-back of the metric  $g$  to the world-line of the particle,  $\mathcal{P}$ , embedded in  $M$  via  $X : \mathcal{P} \hookrightarrow M$ . Show that the equation of motion of (4.8) is given by

$$\frac{d^2 X^\mu}{ds^2} + \Gamma_{\nu\rho}^\mu \frac{dX^\nu}{ds} \frac{dX^\rho}{ds} = 0, \quad (4.9)$$

where

$$ds = \sqrt{-\dot{X}^\mu \dot{X}^\nu g_{\mu\nu}} d\tau, \quad (4.10)$$

and  $\Gamma_{\nu\rho}^\mu$  are the Christoffel symbols given explicitly by

$$\Gamma_{\nu\rho}^\mu = \frac{1}{2} g^{\mu\sigma} (\partial_\nu g_{\rho\sigma} + \partial_\rho g_{\nu\sigma} - \partial_\sigma g_{\nu\rho}). \quad (4.11)$$

Interpret this equation of motion.

(g) The action (4.8) can be generalised to a 1-dimensional string. This leads to the Nambu-Goto action:

$$S = -T \int_{\Sigma} d^2\sigma \sqrt{-\det(X^*g)}, \quad (4.12)$$

where  $T$  is the tension of the string,  $X^*g$  is now the pull-back of  $g$  to  $\Sigma$ , the worldsheet of the string, and  $\sigma^\alpha = (\tau, \sigma)$  denote local coordinates on the worldsheet of the string. Write out (4.12) explicitly.